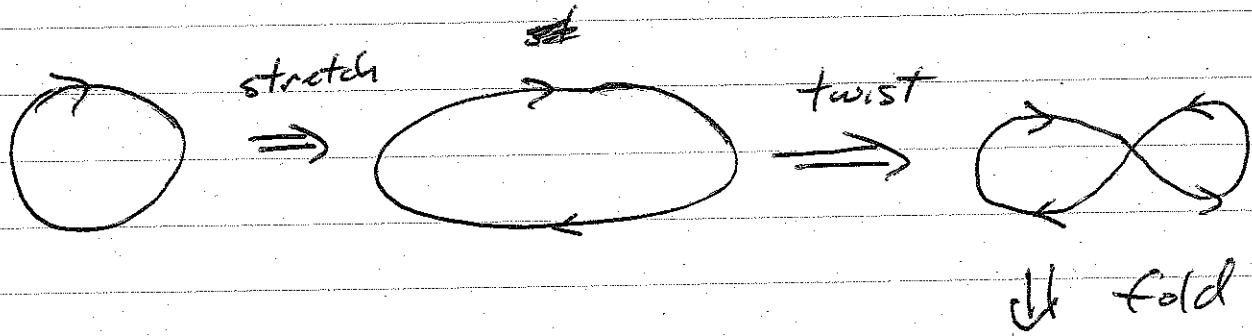


~~Dynamo~~

Generation of Magnetic Fields: the dynamo

Planets, stars, and other objects self-generate magnetic fields, which ~~increase over time~~

The generation is believed to result from the convection of magnetic fields. A simple picture of the amplification process is the stretch-twist-fold dynamo:

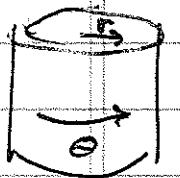


This doubles the poloidal flux. The energy to do this comes from the stretching process.

Cowling's Theorem: no self-generation in a 2-D system

Can a system with 2-D variation amplify B ?

To prove this we consider an axi-symmetric system



\Rightarrow cylindrical geometry

$$\text{with } \frac{\partial}{\partial \theta} C = 0$$

$$B(r, \theta) = B_p + B_\theta \quad u(r, z) = u_p + u_\theta$$

poloidal toroidal

$$\frac{1}{c} \frac{\partial B}{\partial t} + \nabla \times E = 0 \quad \nabla \times B = \frac{4\pi}{c} J$$

$$E = -\frac{1}{c} u \times B + \frac{4\pi}{c} J$$

Want to separate the induction eqn into poloidal and toroidal components

Toroidal:

$$\text{Take } \nabla \theta \cdot \left[\frac{\partial B}{\partial t} + \dots \right] = 0$$

$$\frac{1}{c} \frac{\partial}{\partial t} \nabla \theta \cdot B + \underbrace{\nabla \theta \cdot \nabla \times E}_{-\nabla \theta (\nabla \theta \times E)} = 0$$

$$\frac{1}{c} \frac{\partial}{\partial t} \nabla \theta \cdot B + \nabla \cdot \underbrace{\frac{1}{c} \nabla \theta \times (u \times B)}_{\nabla \theta \cdot B u - \nabla \theta \cdot u B} - \nabla \cdot (\nabla \theta \times \frac{4\pi}{c} J) = 0$$

$$\frac{2}{c} \frac{\partial}{\partial t} \nabla \theta \cdot B + u \cdot \nabla \left(B \cdot \nabla \theta \right) - B \cdot \nabla u \cdot \nabla \theta$$

$$+ 2 \nabla \theta \cdot \nabla \times \frac{4\pi}{c} J$$

$$\underbrace{\frac{1}{c} \nabla^2 r B_\theta}_{-\frac{3C^2}{4\alpha} \nabla \theta \cdot \nabla^2 B}$$

$$- \frac{3C^2}{4\alpha} \nabla \theta \cdot \nabla^2 B = - \frac{3C^2}{4\pi} \frac{1}{r} \left(\nabla^2 B_\theta - \frac{1}{r^2} B_\theta \right)$$

$$\nabla^{\star 2} = \frac{\partial^2}{\partial r^2} + r \frac{1}{r} \frac{\partial}{\partial r} - \frac{1}{r^2}$$

$$\frac{\partial}{\partial r} \frac{B_\theta}{r} + u_p \cdot \nabla \left(\frac{B_\theta}{r} \right) - B_p \cdot \nabla \frac{u_p}{r}$$

$$- \frac{3c^2}{4\alpha} \frac{1}{r^2} \nabla^{\star 2} r B_\theta = 0$$

Polooidal: In a cylinder with $\frac{\partial}{\partial \theta} = 0$ only term that produces a poloidal component is E_θ

$$\frac{\partial}{\partial r} B_p = \nabla \times (u_p \times B_p) = q \nabla \times J_\theta$$

Can write $B_p = \nabla \psi \times \nabla \theta$ since $\nabla \cdot B_\theta = 0$
 $\nabla \cdot B_p = 0$

$$\Rightarrow B_p = \nabla \times \psi \nabla \theta$$

$$B_p \equiv G \times \nabla \theta$$

$$\begin{aligned} \nabla \cdot B_p &= 0 = \nabla \cdot G \times \nabla \theta \\ &= \nabla \theta \cdot \nabla \times G \\ &\Rightarrow G = \nabla \psi \end{aligned}$$

$$\frac{\partial}{\partial r} \psi \nabla \theta + \underbrace{u_p \times B_p}_{u_p \times (\nabla \psi \times \nabla \theta)} + \underbrace{3c J_\theta}_{(-u_p \cdot \nabla \psi) \nabla \theta} + \cancel{\nabla \Phi} = 0$$

no $\nabla \theta$ component

~~$$\begin{aligned} u_p J_\theta &= \nabla \times (\nabla \psi \times \nabla \theta) \cdot \hat{J}_\theta = -B^2 \psi \nabla \theta \cdot \nabla \psi \cdot \nabla \theta \\ &= -\frac{1}{2} \left[\frac{1}{r^2} + \nabla^2 \psi - \frac{\partial^2 \psi}{\partial r^2} \right] \hat{J}_\theta \end{aligned}$$~~

$$\frac{4\pi}{c} \nabla \theta \cdot \vec{J} = \nabla \theta \cdot [\nabla \times (\nabla \theta \times \nabla \phi)] \\ = - \nabla \cdot [\nabla \theta \times (\nabla \times \nabla \phi)] \\ = - \nabla \cdot [\nabla \theta]^2 \nabla \phi$$

$$\frac{4\pi}{c} J_\theta = -r \nabla \cdot \frac{1}{r^2} \nabla \phi$$

$$\frac{\partial \phi}{\partial r} + \mu_p \cdot \nabla \phi - \underbrace{\frac{3c^2}{4\pi} r^2 \nabla \cdot \frac{1}{r^2} \nabla \phi}_{\nabla^2 \phi} = 0$$

$$\boxed{\frac{\partial \phi}{\partial r} + \mu_p \cdot \nabla \phi - \frac{3c^2}{4\pi} \nabla^2 \phi = 0}$$

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$$\frac{\partial^2 \theta}{\partial r^2} + \frac{1}{r} \frac{\partial \theta}{\partial r} - \frac{3C^2}{4\alpha} \left[\frac{\partial^2 \theta}{\partial t^2} + \frac{\partial^2 \theta}{\partial x^2} \right] = 0$$

\Rightarrow convective/diffusion equation

\Rightarrow no coupling to B_θ

\Rightarrow no source

$\Rightarrow B_p$ does away.

\Rightarrow source due to B_p in B_θ eqn $\rightarrow 0$

$$\left(\frac{\partial}{\partial t} + U_p \cdot \nabla \right) \frac{B_\theta}{r} - 3 \frac{1}{r^2} \nabla^2 r B_\theta = 0$$

\Rightarrow again convection

$\Rightarrow B_\theta \rightarrow 0$

Need $B_p \rightarrow B_\theta \rightarrow B_p \Rightarrow$ closed loop.

Note that $\mathcal{R} = \frac{U_p}{\Omega}$ is rotation rate, so that

If \mathcal{R} varies along B_p the rotation will twist B_p into the B direction

\Rightarrow not feed back on B_p so once $B_p \rightarrow 0$ this ends

The $\alpha \beta \gamma$ kinematic dynamo

Now consider the 3-D case. To make progress assume that the flows and magnetic field can be separated into small scale and large scale components

$$\vec{B} = \vec{B}_0 + \vec{\tilde{B}} \quad \text{# take } \vec{B} \text{ to be small}$$

$$\vec{v} = \vec{v}_0 + \vec{\tilde{v}} \quad \langle \tilde{v} \rangle = 0, \langle \tilde{B} \rangle = 0$$

\Rightarrow take $\vec{\tilde{B}}, \vec{\tilde{v}}$ to be small.

\Rightarrow write down equations for \vec{B}_0, \vec{v}_0

$$\frac{d}{dt} \vec{B}_0 = \nabla \times (\vec{v}_0 \times \vec{B}_0) - \nabla \times \vec{\tilde{v}}_0 + 3 \nabla^2 \vec{B}_0$$

$$\frac{d}{dt} \vec{v}_0 = \nabla \times (\vec{v}_0 \times \vec{B}_0 + \vec{\tilde{v}}_0 \times \vec{B}_0) + \nabla \times \vec{\tilde{v}}_0 + 3 \nabla^2 \vec{v}_0$$

$$\vec{\tilde{v}}_0 = -\langle \vec{v} \times \vec{\tilde{B}} \rangle = \text{average electric field from turbulence.}$$

$$\vec{\tilde{v}}_0 = \vec{v} \times \vec{\tilde{B}} - \langle \vec{v} \times \vec{B} \rangle$$

\Rightarrow neglect $\vec{\tilde{v}}_0$ since small compared with other terms

\Rightarrow keep $\vec{\tilde{v}}$ because only coupling term for generating B_0 .

To evaluate $\bar{\epsilon}$ need to find the correlation between \tilde{v} and \tilde{B} . This is obtained from the \tilde{B} equation \Rightarrow specifically the $\nabla \times (\tilde{v} \times \underline{B}_0)$ drive term.

Consider a local region where

Jump to a local frame where $\underline{v}_0 \approx 0$

$$\tilde{\underline{B}} = \int dt' \nabla \times (\tilde{\underline{v}}(x, t') \times \underline{B}_0),$$

where the time dependence of \underline{B}_0 is neglected

$$\bar{\epsilon}_m = - \int dt' \langle \tilde{\underline{v}}(x, t') \times (\nabla \times (\tilde{\underline{v}}(x, t') \times \underline{B}_0)) \rangle$$

If the turbulence is isotropic

~~of the same length scale~~ there are only

two possible vectors which can define

the direction for $\bar{\epsilon}_m = \underline{B}_0$ and $\nabla \underline{B}_0$

Thus,

$$\bar{\epsilon}_m = \alpha \underline{B}_0 + \beta \nabla \times \underline{B}_0$$

α portion \Rightarrow take \underline{B}_0 constant along c axis

$$\bar{\epsilon}_{\alpha} = - \int dt' \langle \tilde{\underline{v}}_m \times \underline{B}_0 \cdot \nabla \underline{v}_m \rangle$$

only c_i survives because it is along

$$\bar{\epsilon}_{\alpha i} = - \int dt' \langle \underline{B}_0 \cdot (\tilde{\underline{v}}_i \times \nabla \underline{v}'_i) \rangle$$

$$= - \underline{B}_0 \cdot \left(\int dt' \langle \tilde{v}_j \tau_i \tilde{v}'_k - v_k \tau_i v_j \rangle \right)$$

2 portion

$$\mathbb{E}_n \alpha = - \int dt' \langle \tilde{v} \times B_0 \cdot \nabla \tilde{v}' \rangle$$

take B_0 to be locally in the i direction

$$\mathbb{E}_n \alpha = - \int dt' B_{0i} \langle \tilde{v} \times \nabla_i \tilde{v}' \rangle$$

for isotropic \tilde{v}

$\langle \tilde{v} \times \nabla_i \tilde{v}' \rangle$ will be on average zero unless it is in the ∇_i direction

\Rightarrow

$$\mathbb{E}_{n,i} \alpha = - \int dt' B_{0i} \underbrace{\langle \tilde{v} \times \nabla_i \tilde{v}' \rangle}_{\langle \tilde{v}_j \nabla_i \tilde{v}'_k - \tilde{v}_k \nabla_i \tilde{v}'_j \rangle}$$

isotropy means can cycle indices



$$\langle \tilde{v}_k \tilde{v}_i \tilde{v}'_j \rangle = \langle \tilde{v}_j \tilde{v}_k \tilde{v}'_i \rangle$$

$$\begin{aligned} \Sigma_{\alpha i} &= + B_{0i} \left\langle \tilde{v}_j (\nabla \times \tilde{v}')_j \right\rangle dt' \\ &= \frac{1}{3} B_{0i} \left\langle \tilde{v} \cdot \nabla \times \tilde{v}' \right\rangle dt' \end{aligned}$$

$$\alpha = \frac{1}{3} \int^t dt' \left\langle \tilde{v} \cdot \nabla \times \tilde{v}' \right\rangle$$

$$= \frac{1}{3} \int^t dt' \left\langle \tilde{v}_m \cdot \tilde{\omega}'_m \right\rangle$$

$$\boxed{\alpha = \frac{1}{3} \tau \langle \tilde{v}_m \cdot \tilde{\omega} \rangle}$$

may have either
sign - kinetic
helicity

τ = coagulation time

$$\Sigma_{\beta} = + \int^t dt' \left\langle \tilde{v} \times \cancel{(\tilde{v}' \cdot \nabla)} B_0 \right\rangle$$

$$\Rightarrow \text{only } \langle \tilde{v}_i \tilde{v}'_j \rangle = 0 \text{ for } i \neq j$$

$$\begin{aligned} \Sigma_{\beta} &= \int^t dt' \left\langle E_{ijk} \hat{e}_i \tilde{v}_j \tilde{v}'_k \nabla_p B_0 \right\rangle \\ &\quad \left\langle E_{ijk} \hat{e}_i \tilde{v}_j \tilde{v}'_k \nabla'_j B_0 \right\rangle \end{aligned}$$

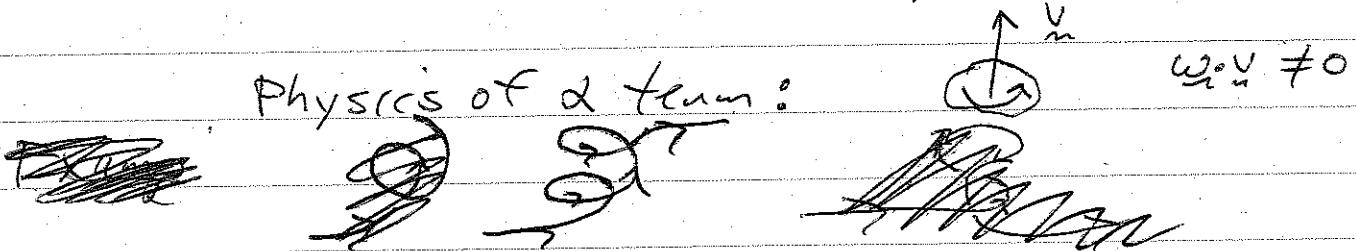
$$\frac{1}{3} \left\langle E_{ijk} \hat{e}_i \nabla_j B_0 \right\rangle \langle \tilde{v}_i \cdot \tilde{v}'_k \rangle$$

$$\frac{1}{3} \langle \tilde{v} \cdot \tilde{v}' \rangle \nabla \times B_0$$

$$\beta = \frac{1}{3} \int^t dt' \left\langle \tilde{v}_i \cdot \tilde{v}'_i \right\rangle = \boxed{\frac{1}{3} \tau \langle \tilde{v}_i \cdot \tilde{v}'_i \rangle}$$

$$\frac{\partial}{\partial t} \mathbf{B}_0 = \nabla \times (\mathbf{V}_0 \times \mathbf{B}_0) - \nabla \times \alpha \mathbf{B}_0 + (\beta + \gamma) \nabla^2 \mathbf{B}_0$$

$\Rightarrow \beta$ produces an enhanced flux diffusion
 \Rightarrow anomalous resistivity



Flux Egn As before $\frac{\partial}{\partial t} \nabla \times \mathbf{B}_0 + (\gamma) = - \nabla \times \alpha \mathbf{r} \mathbf{B}_0$

$$\frac{d}{dt} \mathcal{A} = \gamma \nabla^2 \mathcal{A} + \alpha r \mathbf{B}_0$$

\Rightarrow azimuthal field now generates poloidal field so Ampere's theorem fails

example

\Rightarrow neglect \mathbf{V}_0 , $\alpha = \text{const}$

$$\frac{\partial}{\partial t} \mathbf{B}_0 = - \nabla \times \mathbf{r} \mathbf{B}_0 \quad \text{at } \mathbf{B}_0 \sim e$$

$$\frac{\partial}{\partial t} \nabla \times \mathbf{B}_0 = \alpha \nabla^2 \mathbf{B}_0$$

$$\mathbf{r} \mathbf{B}_0 = - \alpha \frac{1}{8} \alpha (-k^2) \mathbf{B}_0$$

$$\boxed{\alpha^2 = \omega^2 k^2}$$

\Rightarrow amplification of field